1	Revision - GRL 07/27/2022
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3	Seasonal Variations of Medium-Scale Waves observed by ICON-MIGHTI
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17	Highlight
18	1. Medium-Scale Perturbations (MSP) are identified in MIGHTI temperature and wind profiles
19	from 90 to 250 km in altitude.
20	2. MSP show semi-annual variations below ~130 km and annual variations above ~160 km.
21	3. Seasonal variations of MSP are influenced by background winds and tropospheric sources.
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24 Abstract

25 This work presents an analysis of seasonal variations of medium-scale perturbations (~500 26 to ~5700 km) spanning altitudes from 90 km to 250 km using temperature and wind measurements 27 made by the Michelson Interferometer for Global High-resolution Thermospheric Imaging (MIGHTI) instrument onboard the Ionospheric Connection Explorer (ICON) in the latitude range 28 29 of 0°-40° N during 2020-2021. Both Medium-scale perturbations (MSP) in temperature and winds below ~120 km show semi-annual variations, whereas annual variations of MSP for winds become 30 dominant between 160 km to 250 km. The largest wind MSP was observed at ~110-120 km 31 32 throughout the year. Spatial variations of MSP at 90-250 km do not show clear geographic patterns 33 in either temperature or wind. Our analysis suggests both seasonal variations of MSP between 90 34 and 250 km altitudes are influenced by variation on both the sources of MSP and changes in the 35 background wind.

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37 1. Introduction

Gravity waves are small-scale atmospheric waves that distribute momentum and energy 38 39 from source regions over large geographic distances and altitude and exert significant influence 40 on the dynamics of the atmosphere, from the troposphere to the thermosphere and the ionosphere [e.g., Fritts and Alexander, 2003, Vadas and Liu, 2009; Yigit et al., 2009; Miyoshi et al., 2014; 41 42 Yiğit, 2017; Yiğit and Medvedev, 2017, Becker and Vadas, 2018]. Vadas and Liu [2013] and Liu 43 and Vadas [2013] presented that convectively generated gravity waves propagate to the 44 thermosphere and their wave breaking have significant influences on the ionosphere and the 45 thermosphere using convective model, ray-tracing, and TIME-GCM model. Yigit et al. [e.g., 2009, 46 2012, 2017] extended GW parametrization up to the thermosphere and presented a strong GW

47 influence on the thermosphere and the ionosphere. Their parametrization results are supported by high-resolution simulations of Miyoshi et al. [2015]. Miyoshi et al. [2015] uses a whole 48 49 atmospheric model (GAIA model) that resolves GW from the lower atmosphere to the 50 thermosphere (~500 km) and confirmed strong GW drag in the thermosphere. In addition to 51 propagations from the lower atmosphere, secondary/tertiary generation of gravity waves by 52 dissipation of primary gravity waves that are generated in the troposphere and the stratosphere can 53 be an important source in the thermosphere [e.g., Vadas, 2013; Becker & Vadas, 2018]. These previous modeling studies have shown the importance of GWs on the variabilities in the 54 55 ionosphere and the thermosphere.

56 Observations of gravity waves in the thermosphere have been reported using GOCE 57 satellite observations [e.g., Forbes et al., 2016; Liu et al., 2017]. Liu et al. [2017] presented annual 58 and semi-annual gravity wave variations with the peak in June-August using GOCE neutral density between 220 to 280 km altitudes (average height is ~250 km). Forbes et al. [2016] showed 59 60 hemispheric differences and latitudinal variations of GOCE gravity wave density variance. Besides GOCE gravity wave studies, thermospheric global gravity wave observations are rare. In 61 62 particular, there is a lack of global gravity wave observations between the MLT (the mesosphere 63 and the lower thermosphere) and ~ 250 km altitude region. Additional observational studies to 64 understand distributions of medium-scale perturbations that includes gravity waves can be 65 beneficial for validations of whole atmospheric models in the thermosphere. Since December 2019, 66 the Michelson Interferometer for Global High-Resolution Thermospheric Imaging (MIGHTI) 67 instrument on the Ionospheric Connection Explorer (ICON) satellite has been providing winds from 90 km to 290 km and temperature measurements from 90 to 127 km in daytime, with a more 68 69 limited altitude range at night. In this work, using MIGHTI-temperature and winds, perturbations

70 with horizontal scales smaller than ~5000-6000 km are extracted and defined as medium-scale 71 perturbations (MSP), which include gravity wave perturbations. Results of temporal, spatial, and 72 altitudinal variations of medium-scale wave variances are shown and discussed in this work and 73 provide new insight into medium-scale perturbation variation in the mesosphere and the 74 thermosphere.

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2. ICON-MIGHTI and Analysis Method

The Ionospheric Connection Explorer (ICON) was launched October 10, 2019 [Immel et 77 al., 2018]. Temperatures and winds in the upper atmosphere have been measured by the Michelson 78 Interferometer for Global High-Resolution Thermospheric Imaging (MIGHTI) instrument [Immel 79 80 et al., 2018; Englert et al. 2017; Harlander et al., 2017; Harding et al., 2017; Stevens et al., 2018]. 81 In this study, MIGHTI-A temperatures (version 5) from 90 km to 127 km, MIGHTI Greenline winds (version 4) from ~95 km to 150 km and MIGHTI Redline winds (version 4) from 160 km 82 83 to 250 km are used. Data are continuous through these altitude ranges only in daytime so we focus on the 10 to 16 hour local time range (LT hereafter). The vertical sampling of MIGHTI temperature 84 85 vertical resolution is ~3 km. The vertical sampling frequency of MIGHTI green-line wind is also 86 \sim 3 km, and is \sim 10 km for red-line wind. MIGHTI redline winds and greenline winds have been 87 validated against ground based FPI and radar observations, respectively [Makela et al., 2021] 88 Harding et al., 2021]. Validations of version 4 MIGHTI temperature against ground-based lidar 89 observations can be found in Yuan et al. [2021].

90 Characteristics of MSP are determined from these data by first subtracting background and
91 large-scale waves from the observations. This method is widely used for gravity wave analysis
92 using various satellite observations including SABER, COSMIC, and HIRDLS [e.g., Ern *et*

93 al., 2011; Preusse et al., 2009; Yamashita et al., 2013; Cullens et al., 2015]. The analysis method is briefly summarized in the following. For each day of data, MIGHTI winds and temperature data 94 95 are binned/averaged into 30-minute LT samples. Background and large-scale waves are then 96 estimated by fitting wavenumber 0 to 6 at fixed LT. By fitting wavenumber 0-6 at fixed LT, both zonal-mean backgrounds, tides, and stationary planetary waves are included, even though none of 97 98 these components can be specifically identified. Estimated background temperature and wind are subtracted from each temperature and wind profile. Residuals (i.e., the difference between the data 99 and the large-scale wave fit) are analyzed to characterize the remaining medium-scale 100 101 perturbations (MSP), which includes gravity wave perturbations. MSP shown in this work 102 characterize waves with horizontal wavelengths longer than ~500 km and smaller than ~4300-5700 km. The analysis is conducted using data from 0° - 40° N because of large numbers of missing 103 104 data points due to South Atlantic Anomaly (SAA) that cause zonal coverage to be incomplete in the southern hemisphere and therefore difficult to fit for large-scale waves. 105

Figure 1 shows an example of MSP analysis. Figure 1a, 1c, and 1e show MIGHTI wind and temperature observations (black dots) and estimated background (combined wavenumber 0-6). Differences between estimated background (redline) and original measurements (black dots) are considered to be medium-scale perturbations in this work. Figures 1b, 1d, and 1f show vertical profiles of MSP with wave-like perturbations for (b,d) zonal wind and (f) temperature. Red dashed lines indicate uncertainties for wind measurements in Figure 1b and 1d and temperature measurements in Figure 1f.



Figure 1. The red line indicates the background zonal winds (wavenumber 0-6) and the black
dots indicate MIGHTI measurements for (a, c) the zonal winds measurements at 213 km and at
108 km, respectively, and (e) temperature measurements at 100 km at 14 LT. The vertical
profiles of medium-scale perturbations for (b) red line zonal winds, (d) green line zonal winds,
and (f) temperature are shown in black lines, and red dotted lines indicate uncertainties of wind
and temperature measurements.

3. Results

123 **3.1.** Climatology of Medium-Scale Perturbations

124 Figure 2 shows the seasonal and latitudinal variations of MSP. The wind MSP variances 125 at 200 km and 110 km are depicted in Figures 2a and 2c, respectively, and temperature MSP at 126 110 km and 94 km are illustrated in Figures 2b and 2d, respectively. MSPs are separated into 127 different latitude ranges of 0-15°N, 15-30°N, and 30-40°N. MSP at 200 km in Figure 2a have the largest variance in June-July at all latitude range (0°-15°N, 15°-30°N, and 30°-40°N). There are 128 secondary peaks in December- January. Peaks MSP variance at 200 km at 15°-30°N and 30°-40°N 129 in June and July are 85-95.0 m²/s² and secondary peak in December and January are $\sim 60 \text{ m}^2/\text{s}^2$. 130 131 The June and July peaks are ~17-26% larger than MSP in December and January in the latitude range of 15°-30°N and 30°-40°N. In the latitude range of 0°-15°N, MSP in June-July are weaker 132 133 and result in weaker annual variations but larger semi-annual variations (in particular for 2021 results). Minimum MSP for the latitude range of 0°-15°N in March-April and September-October 134 135 are $\sim 30 - 40 \text{ m}^2/\text{s}^2$, which is $\sim 40-50\%$ lower than the peaks in June-July. Although there are MSP 136 peaks in December – January, the annual-variations of MSP is generally dominant at 200 km.

137 In contrast to MSP variances at 200 km, MSPs at 110 km at all latitudes indicate a dominant 138 semi-annual variation at the two peaks in December-January and June-August. For the latitude range of 15°-30°N, the peak variance in July is \sim 250-270 m²/s² and December-January peak is 139 140 ~250 m²/s². Semi-annual variations are largest at 15°-30°N. Minimum MSP in Septembers in the 141 year of 2020 and 2021 are ~180 m²/s² compared to maximum ~260 m²/s² in June-July at 15°-30°N. For the latitude range of 0°-15°N, both July and January peaks are $\sim 200 - 250 \text{ m}^2/\text{s}^2$. Forbes et al. 142 143 [2015] showed GOCE gravity wave variations in the altitude range of ~220 to 280 km. Their 144 results also show the largest peak in June-July and dominate annual variations along with



secondary peak in December-January in low latitudes. Our seasonal variations are comparable to

previous work at ~250 km.





Figure 2. Seasonal variations of monthly-zonal-mean MSP variances in (a) zonal-winds at 200
km, (b) temperature at 110 km, (c) zonal winds at 110 km, (d) temperature at 94 km averaged
over (black) 0-15°N, (blue) 15-30°N, (red) 30-40°N from January 2020 to December 2021.
Error bars represent one sigma.

154 Temperature MSP variances at 110 km also show a clear semi-annual variation in all the latitude range with peaks in June-July and November-February in Figure 2b. Peak MSP variances 155 in June-July and November-February are ~100-120 K² and minimum MSP in March and 156 157 September are $\sim 60-70 \text{ K}^2$. MSP variance in solstice is $\sim 30-40\%$ larger than those in equinox. At 158 94 km in Figure 2d, temperature MSP variances at 15°-30°N and 30°-40°N both show semi-annual 159 variations with peaks in June-July and December-January as wind MSP variances. Interestingly, MSPs at 94 km show a dominate annual variation in the latitude range of 0°-15°N, which is 160 different variations from other latitudes. Gravity wave measurements from SABER temperature 161 162 observations at 95 km also found semi-annual variations with two peaks around June-August and December-January averaged over 12°S - 12°N and from 2002 to 2006 [Preusse et al., 2009]. A 163 164 cause of differences in seasonal variations between MIGHTI and SABER could be due to the 165 difference in LT coverage. Preusse et al. [2009] showed SABER observations covering all LT, but in this study MIGHTI observations in Figure 2 focus only on 10-16 LT to cover wide altitude range 166 167 (night time temperature observations only go up to 105 km). When we include all local time for MIGHTI below 105 km and recalculate temperature MSP, a semi-annual variation begins to show 168 169 (Figure not shown). In addition, SABER (5-year average) and MIGHTI (2 years of analysis) 170 observational periods are different. Further comparison between two datasets will be conducted in the future work. 171

To further study altitudinal variations of MSPs, Figure 3 shows the height variation of the monthly mean MSP variance from January 2020 to December 2021. MSP variances are averaged over 0°-40°N, which include all MSP shown in Figure 2. Figures 3a and 3b show seasonal variations of MSP from MIGHTI winds. Wind MSPs are separated in Figure 3a and 3b to indicate winds derived from red-line and green-line emission, respectively. The MSP variances (94 – 150 km altitudes) in Figure 3b show clear semi-annual variations with peaks in December-January and
June-August around solstice. MSP variance at 110-120 km exhibits the largest amplitudes. MSP
in June-July in 2021 are weaker than those in 2020 in Figure 3b about ~5-10%.

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Figure 3. Seasonal variations of monthly-mean MSP variances from (a) MIGHTI red-line zonal
winds (m²/s²), (b) MIGHTI green-line zonal winds (m²/s²), (c) MIGHTI temperature (K²)
averaged over 0°-40°N. The vertical black line marks the end of 2020.

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Figure 3a shows MSP variances from 160 km to 250 km. As shown in Figure 2a, MSPabove 180 km have strong annual variations with a peak in June-July. These peaks are evident in

both 2020 and 2021. From 160 km to 250 km, MSP amplitudes are decreasing with altitude, which
may indicate dissipation of gravity waves propagating from the lower atmosphere [Fritts and
Alexander, 2003]. Figure 3c shows temperature MSP variations from 90 km to 125 km. Semiannual variations are dominated and temperature MSPs maximize between ~105 km and 120 km
altitudes, which is consistent with wind MSP observations shown in Figure 3b.

193 Longitudinal variations of temperature and wind MSP are examined; however, there is no clear longitudinal variations of MSP amplitudes in all altitudes. Miyoshi et al. [2014] presented 194 resolved gravity waves in whole atmospheric models. Their results indicate strong connections of 195 196 gravity wave in the troposphere to gravity waves in the thermosphere around the region of Andes 197 Mountains but not as clear in the low latitude [Miyoshi et al., 2014]. MIGHTI only observes in the 198 latitude range of 10°S-40°N and does not include Andes Mountain area. GOCE satellite results 199 also did not find a geographically located source of strong gravity wave activities, though GOCE found significant latitudinal variations [Forbes et al., 2016]. Preusse et al. [2009] used ray-tracing 200 201 of gravity waves to explain SABER gravity wave variations at ~95 km. Their results indicate that 202 longitudinal variations of gravity waves in the lower atmosphere disappear as they propagate to 95 203 km because gravity waves spread out as they propagate upward. At MIGHTI observed altitudes 204 (90-250 km), spatial variations of monthly mean gravity waves seem not to be strongly tied to tropospheric source distributions, which is consistent with previous satellite gravity wave 205 206 observations.

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208 **4. Discussion**

In this section, the conditions of sources and propagation of gravity waves will be discussed
to explain our observed MSP variations. Figure 4a shows MIGHTI zonal and monthly mean zonal

211 winds from 94 km to 280 km averaged over 10 to 16 LT as in the analyses of Section 3. Zonal 212 mean winds at the altitude range of ~110-150 km show clear annual variations with strong 213 westward winds from May to September. The zonal winds above ~160 km are westward winds 214 throughout the year with largest westward winds in May-August. In order to understand gravity 215 wave propagations below ~95 km, zonal winds from Specified Dynamics Whole Atmosphere 216 Community Climate Model with thermosphere and ionosphere extension (SD-WACCMX) are 217 used in this study and shown in Figure 4b. SD-WACCMX is nudged by GEOS5 below ~60 km to simulate a realistic lower atmosphere [e.g., Kunz et al. 2011; Smith et al., 2017]. To validate 218 219 SD-WACCMX zonal wind above 94 km against MIGHTI zonal wind, SD-WACCM-X data is 220 also averaged over 10 LT to 16 LT and in the latitude range of 0°-40°N. Coincident MIGHTI wind 221 altitudes are indicated by a red box in Figure 4b. Between ~ 90 km to ~ 125 km, SD-WACCM-X 222 show eastward winds consistently through the year. Although MIGHTI shows westward winds around August and October-November, MIGHTI show no clear seasonal variations with mostly 223 224 eastward winds from 95 km to ~115 km as similar to SD-WACCM-X. Between 125 km and 300 225 km, SD-WACCM-X show annual variations with westward from May to August and 226 eastward/weak westward winds from September to ~March. On the other hand, MIGHTI zonal 227 winds are westward through the year between 160 and 280 km. Differences between measured 228 winds and SD-WACCM-X could be due to the lack of gravity wave drag in these heights in SD-229 WACCM-X. Because MIGHTI zonal wind only includes daytime data, differences between 230 MIGHTI and SD-WACCMX could also be coming from differences in tidal amplitudes and/or 231 phases. Although seasonal variations are not consistent between SD-WACCM-X and MIGHTI 232 from ~160 km to 250 km, magnitudes of the zonal winds in July-August are ~60-70 m/s and 233 consistent between SD-WACCM-X and MIGHTI.

Preusse et al. [2009] show latitudinal variations of gravity waves using SABER 234 235 observations and a ray-tracing model from 30 km to 90 km. They showed that gravity waves in 236 the tropics in the troposphere and stratosphere spread out to wide range of latitudes range as they 237 propagate upward to the mesosphere and the lower thermosphere. Therefore, MIGHTI MSP in the 238 latitude range of 0°-40°N at the MLT region could be coming from convective region (20°S-20°N) 239 in the lower atmosphere. To understand wave propagations, Figure 4c shows zonal mean zonal wind from 10 km to 100 km from SD-WACCMX averaged over the tropics (latitude range of 240 241 20°S-20°N). Zonal winds in Figure 4c are daily mean results including all local time. SD-242 WACCMX zonal winds show clear semi-annual variations, which is consistent with previous 243 studies of semi-annual oscillations (SAO) [e.g., Garcia et al., 1997; Smith et al., 2017]. The periods 244 of westward phases of SAO in WACCM are shorter than those in HRDI observations [Garcia et al., 1997]; however, the general structures of SAO in WACCMX are compared well with Garcia 245 et al. [1997]. Such semi-annual variations in zonal wind structure can be a cause of observed semi-246 247 annul variations of gravity waves in the mesosphere and thermosphere as also discussed by 248 previous studies [e.g., Preusse et al., 2009] and could affect our observed MSP.

Furthermore, according to both MIGHTI and SD-WACCM-X in Figures 4a and 4b, the zonal wind direction changes from eastward to westward in the altitude range of 130-160 km. Changes in wind direction are often associated with dissipations of gravity waves [Fritts and Alexander, 2003], and such wave dissipation can generate secondary generation of gravity waves [e.g., Vadas 2013]. Wind MSP variance in Figures 3a and 3b show changes from semi-annual variations below 150 km to annual variations above 160 km. The altitude of MSP changes (130-180 km) coincide with changes of zonal wind directions seen both in MIGHTI and SD-WACCMX. Our results suggest that background wind structure have an important role in observed MSPseasonal variations.

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Figure 4. Seasonal variations of (a) MIGHTI zonal-mean zonal winds from 95-280 km and (b)
SD-WACCMX zonal mean zonal wind from 10-280 km. Zonal winds in Figure 4a and 4b are
averaged over 10-16 LT and latitude range of 0°-40°N. Positive/negative indicate
eastward/westward zonal winds, respectively. Red boxes indicate MIGHTI observational
altitude. (c) SD-WACCMX zonal and daily (all LT) mean zonal wind in the latitude range of 20S20N. (d) NCEP monthly mean precipitation in 2020 (mm/day).

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We further investigate the connection to gravity wave source variations in the troposphere. Major sources of gravity waves are orography, frontal system, and convection. One of the largest gravity wave sources in the low-latitude troposphere is convection [Fritts and Alexander, 2003].

271 In order to quantify changes in convective activities, Figure 4d shows NCEP monthly mean zonal 272 mean precipitation obtained for the year of 2020. White lines in Figure 4d show the ICON 273 observational latitudes from 10°S to 40°N. The largest peak in precipitations is in June-September in the latitude range of 0°-20°N. The second peak is in January in the latitude range of 0°-20°S. 274 275 These two peaks coincident with MIGHTI MSP semi-annual peaks below 150 km in Figures 3b 276 and 3c. Therefore, our observed MIGHTI MSP can be influenced by both semi-annual variations 277 of convective sources in the low latitudes and also zonal wind structures, which consistent with previous studies in the mesosphere region. 278

Furthermore, geomagnetic activity can also be a source for small to medium scale waves. Bruinsma and Forbes [2008] presented the influences of geomagnetic activity (Kp > 4.5) on smallmedium scale waves (600-2400 km) in relative neutral density from high-latitudes to low-latitudes. Although it is possible that geomagnetic activities can affect our observed MSP, there are not many strong geomagnetic activity during our observed time. Influences of geomagnetic activities on medium-scale waves are important and will be further investigated in the following studies.

285

286 **5.** Conclusions

This work presents the first seasonal variations of meidum-scale perturbations from 90 km to 250 km using both temperature and winds in the latitude range of 0°-40°N from January 2020 to December 2021. The MIGHTI instrument onboard the ICON has been providing zonal and meridional winds from 94 km to 280 km and temperature measurements from 90 to 127 km during daytime since December 2019. Medium-scale perturbations (wavelengths between ~500 km and ~5700 km) are extracted from MIGHTI temperature and wind profiles. MSP amplitudes in temperature and winds in the altitude range of ~100-130 km shows clear semi-annual variations, 294 and annual variations of MSP become dominant between 180 km to 250 km. Largest MSP 295 variances are observed at ~110-120 km altitude throughout the year from both winds and 296 temperature. Spatial variations of MSP do not present clear and consistent longitudinal structure 297 in the altitude range of 90-250 km and in the ICON observational range of 0°-40°N in both 298 temperature and winds. This is likely due to changes in background winds below 95 km and that 299 gravity waves originating in the lower atmosphere spread out as they propagate. Our analysis of zonal winds and precipitation data suggests both gravity wave sources and background wind 300 changes can affect seasonal variations of MSP from 90 to 250 km. 301

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311

312 Open Research

313 Data Availability Statement

314 This analysis used version 04 of the Level 2 ICON MIGHTI wind data and version 05 of the ICON 315 MIGHTI temperature which available the ICON website data are from 316 (https://icon.ssl.berkeley.edu/Data) NASA's Space Physics Facility and Data

- 317 (<u>https://cdaweb.gsfc.nasa.gov/pub/data/icon/</u>). SD-WACCMX is available from
 318 <u>https://doi.org/10.5065/rjgt-g951 [Maute, 2022]</u>. NCEP precipitation data is available at
 319 https://psl.noaa.gov/data/gridded/data.cmap.html_[Huffman et al., 1997].
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